



PRESS RELEASE (2026/07/10)

Capturing the cosmic 'drift' before a star is born

Researchers observe the velocity difference between two molecules in a prestellar core, elucidating the process of early star formation

Fukuoka, Japan—Stars like our Sun are formed from the collapse of stellar objects called prestellar cores, cold and dense concentrations of gas and dust held together by gravity. While many questions remain on the exact mechanisms of star formation, thanks to advanced radio telescopes, researchers have been able to garner new insights into the inner workings of infant stars.

Now, publishing in [*Astronomy & Astrophysics*](#), researchers from Kyushu University and Max Planck Institute for Extraterrestrial Physics have, for the first time, detected a phenomenon known as ambipolar diffusion occurring in a prestellar core. This phenomenon results in the weakening of the magnetic support of the core, leading to gravitational collapse to form an infant star called a protostar. These new findings provide further insight into the key processes of early star formation, and by extension how stellar systems like ours are created.

"Prestellar cores are fascinating stellar bodies. They are dense and cold, and a source of a lot of complex chemistry. The cold environment allows for molecules to assemble into more complex ones like precursors of prebiotic organic molecules," explains first author [Doris Arzoumanian](#), an Associate Professor at Kyushu University's [Institute for Advanced Study](#). "One of the questions we are investigating is the role of magnetic fields in star formation. Strong magnetic fields permeate prestellar cores. If that field is too strong, it can delay gravitational collapse and therefore star formation. We wanted to investigate how prestellar cores reduce the strength of their magnetic field."

Using the Institute for Radio Astronomy in the Millimetre Range (IRAM) 30 m telescope the research team turned their sights to L1544, a prestellar core located in the Taurus molecular cloud, one of the nearest star-forming regions to Earth. In molecular clouds, gas is partially ionized, meaning ions are strongly coupled to magnetic fields, while neutral particles interact with the field indirectly through collisions. Studying these molecules is the key to understanding the state of the core's magnetic field.

However, because prestellar cores are so cold, the most common molecular tracers freeze onto dust grains, making them invisible. Therefore, the team had to identify a new set of molecules to trace.

"We selected Diazenylium-d₁ (N₂D⁺), an ion, and para-monodeuterated ammonia (para-NH₂D), a neutral molecule, as our tracers because they are generally located in similar high-density regions within prestellar cores," explains second author [Silvia Spezzano](#), group leader at the [Max Planck Institute for Extraterrestrial Physics](#). "We therefore collected spectral data of the core and modeled the velocity of the two molecules."

"This process is known as ambipolar diffusion. Until now, observing this phenomenon in a prestellar core was a major challenge," continues Arzoumanian. "As ambipolar diffusion continues, the strength of the magnetic field decreases. Eventually, gravity becomes the primary driving force in the core, resulting in its gravitational collapse into a protostar."

“These results were possible thanks to an interdisciplinary collaboration of expert observers and theorists in the fields of gas dynamics, astrochemistry, and dust physics,” concludes Arzoumanian. “Understanding star formation addresses a fundamental question about the origin of life in planetary systems and helps us better understand the universe as a whole.”

For more information about this research, see "Probing the ion-neutral drift velocity towards the L1544 prestellar core. Detection of ambipolar diffusion using N₂D⁺ and para-NH₂D" Doris Arzoumanian, Silvia Spezzano, Tommaso Grassi, Paola Caselli, Yusuke Tsukamoto, Haruka Fukihara, Yoshiaki Misugi, Felipe Alves, Jaime Pineda, Sigurd Jensen, Elena Redaelli, and Alexei Ivlev, Astronomy & Astrophysics, www.aanda.org/10.1051/0004-6361/202658871

Founded in 1911, [Kyushu University](#) is one of Japan's leading research-oriented institutions of higher education, consistently ranking as one of the top ten Japanese universities in the Times Higher Education World University Rankings and the QS World Rankings. Located in Fukuoka, on the island of Kyushu—the most southwestern of Japan's four main islands—Kyushu U sits in a coastal metropolis frequently ranked among the world's most livable cities and historically known as Japan's gateway to Asia. Its multiple campuses are home to around 19,000 students and 8,000 faculty and staff. Through its [VISION 2030](#), Kyushu U will “drive social change with integrative knowledge.” By fusing the spectrum of knowledge, from the humanities and arts to engineering and medical sciences, Kyushu U will strengthen its research in the key areas of decarbonization, medicine and health, and environment and food, to tackle society's most pressing issues.

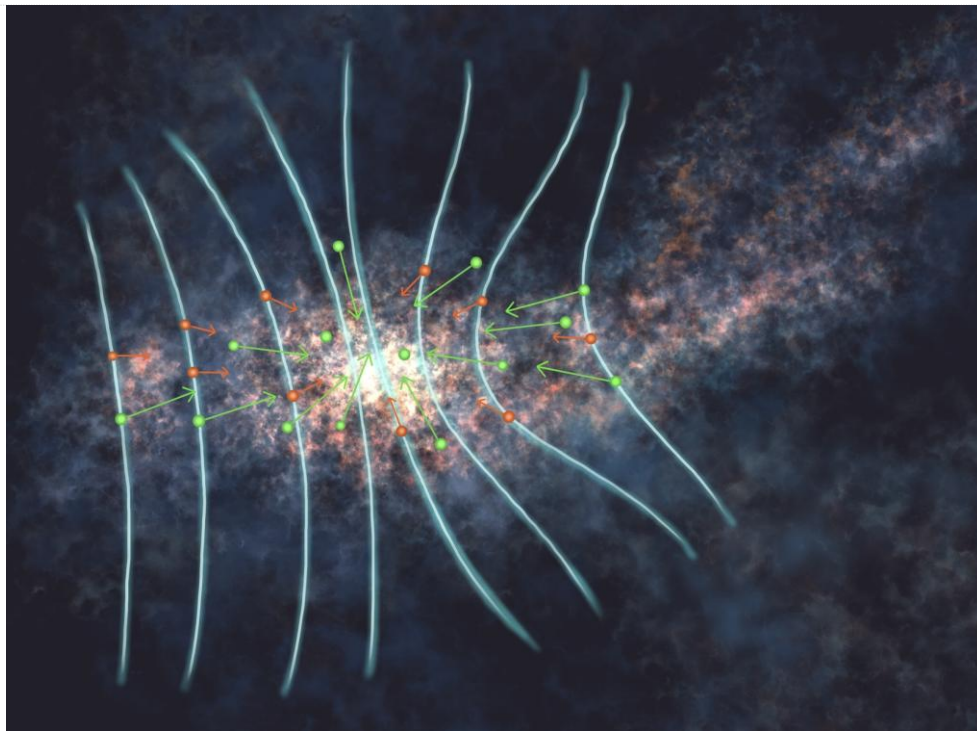


Fig. 1. Illustration of ion-neutral drift in the L1544 prestellar core. The blue lines represent the magnetic field lines, which are bent due to the gravitational contraction of the core. The red and green dots depict the ion and neutral molecular species, respectively, and the arrows trace their inflow motion towards the core center (the faster they travel the longer the arrows). While in the outer part of the core both the ions and neutrals are attached to the magnetic field lines, within the inner part of the core the neutrals decouple from the magnetic field lines and infall faster compared to the ions, which remain attached to the field lines. This ion-neutral decoupling known as ambipolar diffusion is required for the onset of the gravitational collapse of the prestellar core, which will produce a protostar in its center and ultimately a stellar system similar to our own Solar System. (Yurika Nakamura and Doris Arzoumanian/Kyushu University)

[Contact]

Doris Arzoumanian Associate Professor
Institute for Advanced Study

Tel: +81 92-802-4371

E-mail: arzoumanian.doris.958@m.kyushu-u.ac.jp